



LUVOIR/POLLUX First sizing and preliminary thoughts

Workshop, IAP 3/4/17





Introduction

- Aims of this talk
 - To give an overview of the current status of POLLUX
 - Defining the top-level opto-mechanical architecture of POLLUX
 - To share concerns/possible solutions
 - To identify priorities in the requirements (if any)
- All points addressed here are proposals and/or examples... nothing frozen yet !





Top-level Scientific Requirements

- REQ 01: Spectral Resolution: 120 000 (200 000)
- REQ 02: Waveband:
 - REQ 02a: 98-390 nm (90-visible)
 - REQ02b: Minimum order length: 6 nm
 - REQ 02c: Full wavelength range avaiable in one shot
 - REQ 02d: 1-2nm overlap between wavelength range
- REQ 03: SNR
 - SNR=10 for flux 1e-17 erg/s/cm2 in the NV line (124 nm) in a brown dwarf within 40 pc and dispersion 100,000 in 10,000 seconds
- + SNR=10^5 below 156 nm integrating for 15 hours over 0.1 nm a total flux of 5x10^-17 erg cm^-2 s^-1
- REQ 04: Polarization
 - REQ 04a: Circular + linear
 - REQ 04b: Sensitivity: 10^-6 10^-11
- REQ 05: Aperture size: 0.03"
- REQ 06: Observing mode: with and without polarimetry





Top-level Scientific Requirements

- REQ 01: Spectral Resolution: 120 000 (200 000)
- REQ 02: Waveband:

Not considered yet:

Magnitude faintest target: V=26

Radial velocity stability: 1/10th of pixel

Flux stability: 0.001% over 30h

Time resolution: 1sec in FUV and 30sec in NUV

brown r 0.1

- Max exposure time : 30h KEQ 04: Polarization
 - REQ 04a: Circular + linear
 - REQ 04b: Sensitivity: 10^-6 10^-11
- REQ 05: Aperture size: 0.03"
- REQ 06: Observing mode: with and without polarimetry



Echelle Spectrograph Principle

Cnes

ABORATOIRE D'ASTROPHYSIOL







Waveband: 98-390nm

- Need for a spectral separation
 - At least 2 channels are needed (2 octaves)
- Possible (identified) solutions:



=> Limit the amount of channels





Channels

- Proposal:
 - 2+1 channels compatible with optical limits and detector characteristics

FUV	MUV	NUV
• ≈90nm	• ≈120nm	• ≈220nm
• ≈125nm	• ≈220nm	• ≈400nm

- Rationales/drivers:
 - Minimum: 90nm discussed next slide
 - Maximum (400nm): related to the nb of octave
 - Intermediate (220nm): arbitrary (but detector limitations)
 - Overalap: depends on the technology (see architecture)





Minimal WB: 90 nm (goal)



- Telescope Transmission
 - 30%^4 = 8%
- Instrument transmission
 - 50%^3 = 12.5%
- Optical efficiency: 1%
- Without polarization, without QE
- Impact on the transmission in the rest of the band





Example of Possible Architecture







Configuration #1: "fully" compliant



WAVELENGHT RANGE: From 120 to 220 nm Order size (max): 15.644 nm 🗸 Order size (min): 5.989 nm

AVERAGE RESOLUTION: 124585 ✓ Minimum Resolution: 91311 X Maximum Resolution: 208236 ✓ ECHELLE GRATING CHARACTERISTICS: Grating Blaze angle: 63.00 deg ✓ Groove density: 675.00 lines/mm ✗ Not Usual Grating diameter: 66.70 mm by 146.91 mm ✓ Grating orders from 12 to by 22 ✗ Not Usual

CROSS DISPERSER CHARACTERISTICS: Groove density: 450 lines/mm 🖌





Configuration #2: "feasable" (today)



WAVELENGHT RANGE: From 120 to 220 nm Order size (max): 2.688 nm X Order size (min): 0.827 nm X

AVERAGE RESOLUTION: 119875 V Minimum Resolution: 113773 V Maximum Resolution: 126756 V ECHELLE GRATING CHARACTERISTICS: Grating Blaze angle: 63.00 deg ✓ Groove density: 101.00 lines/mm ✓ Grating diameter: 66.70 mm by 146.91 mm ✓ Grating orders from 80 to by 147 ✓

CROSS DISPERSER CHARACTERISTICS: Groove density: 450 lines/mm 🖌





Configuration #3: "realistic" (tomorrow)



WAVELENGHT RANGE: From 120 to 220 nm Order size (max): 7.502 nm Order size (min): 2.551 nm X

AVERAGE RESOLUTION: 124396 V Minimum Resolution: 106571 V Maximum Resolution: 150361 V ECHELLE GRATING CHARACTERISTICS:
Grating Blaze angle: 63.00 deg ✓
Groove density: 304.00 lines/mm ✓
Grating diameter: 66.70 mm by 146.91 mm ✓
Grating orders from 27 to by 49 ✓

CROSS DISPERSER CHARACTERISTICS: Groove density: 450 lines/mm 🖌





Open points impacting the architecture

- Most critical requirements
 - REQ02b: 6 nm/order leading to >600lines/mm
 - REQ02c: all the spectrum in oneshot
- Major identified trade-off:
 - Spectral separation
 - dichroics, mechanism, or any combination of them
 - Polarimeter retractability
 - If confirmed: mechanism + focus compensator to be studied
 - Detector technology and Pixel size:
 - impact of the detector cut-off, characteristics, pixel size, ...
- Feasability studies
 - Echelle Grating with groove density higher than 100lines/mm
 - Coating performance expected
 - impact on the FUV and the need to retract the polarimeter
 - Cross disperser performance (e.g. improved with ion etching)
 - Large focal plane?





Not discussed here (yet)

- Polarimeter
 - Design based on mirrors to meet the shortest wavelength
- Detectors
 - Constraints to be taken into account
- Mass/volume
- And many more ...